

Be stars exhibit changing hydrogen emission features in their spectra, unlike most stars. Using a 14-inch telescope and a spectrograph, calibrated for sensitivity with a well-known stellar spectrum, it is possible to track the relative intensities of these emission lines. I will present results showing variable emission lines of 11 stars over 5 months. Comparison of these features could provide a window into the composition and variability of the circumstellar material producing the emission.